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# The single-particle pseudorelativistic Jansen–Hess operator with magnetic field

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## Abstract

The pseudorelativistic no-pair Jansen–Hess operator is derived for the case where in addition to the Coulomb potential an external magnetic field  $\mathbf{B}$  is permitted. With some restrictions on the vector potential, it is shown that this operator is positive provided the strength  $\gamma$  of the Coulomb potential is below a critical value ( $\gamma_c \leq 0.35$ , depending on the magnetic field energy  $E_f$ ). Moreover, for  $\gamma < 0.32$  and for  $\mathbf{B}$  tending asymptotically to zero in a weak sense, the essential spectrum is given by  $[m, \infty) + E_f$ .

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## 1. Introduction

The spectral properties of the Dirac operator and its nonrelativistic limit, the Pauli operator, describing an atom in an external magnetic field, are a topic of current interest (see the comprehensive review by Erdős [8]). The Dirac operator for an electron in an electric field  $V$  and a magnetic field  $\mathbf{B} = \nabla \times \mathbf{A}$ , acting in the Hilbert space  $L_2(\mathbb{R}^3) \otimes \mathbb{C}^4$ , is given by [3, section 1.3]

$$\begin{aligned} H &= D_A + V + E_f \\ D_A &:= \alpha \mathbf{p}_A + \beta m, \quad \mathbf{p}_A := \mathbf{p} - e\mathbf{A}. \end{aligned} \quad (1.1)$$

$D_A$  is the free Dirac operator with  $\alpha$  and  $\beta$  Dirac matrices,  $m$  is the electron mass,  $V = -\gamma/x$  is the Coulomb field generated by a nucleus of charge  $Z$  fixed at the origin ( $\gamma = Ze^2$  with  $e^2 \approx 1/137.04$  the fine structure constant). In (1.1) the (classical) field energy  $E_f$  is included:

$$E_f := \frac{1}{8\pi} \int_{\mathbb{R}^3} B^2(\mathbf{x}) \, d\mathbf{x} = \frac{1}{8\pi} \|\mathbf{B}\|^2 \quad (1.2)$$

where  $\|\cdot\|$  denotes the  $L_2$ -norm,  $\mathbf{x}$  is the coordinate and  $\mathbf{p} = -i\nabla$  the momentum of the electron. Relativistic units ( $\hbar = c = 1$ ) are used and  $|\mathbf{x}| = x$ . There is a simple relation to the Pauli operator,  $\frac{1}{2m}(\boldsymbol{\sigma}\mathbf{p}_A)^2 = \frac{1}{2m}[(\mathbf{p}_A)^2 - e\boldsymbol{\sigma}\mathbf{B}]$ , where  $\boldsymbol{\sigma}$  is the vector of Pauli spin matrices

[3, section 1.4],

$$D_A^2 = (\mathbf{p} - e\mathbf{A})^2 - e\sigma\mathbf{B} + m^2. \quad (1.3)$$

We need regularity conditions on the vector potential  $\mathbf{A}$  to assure that  $H$  is well defined and self-adjoint. First, we require that

$$\nabla \cdot \mathbf{A} = 0, \quad \|\mathbf{B}\| < \infty. \quad (1.4)$$

These conditions imply the commutation relation  $\mathbf{p}\mathbf{A} = \mathbf{A}\mathbf{p}$  [19, p 438] and  $\mathbf{A} \in L_6(\mathbb{R}^3)$  which results from a Sobolev inequality [9]. The condition  $\mathbf{B} \in L_2(\mathbb{R}^3)$  renders  $E_f$  finite. If, in addition to  $\nabla \cdot \mathbf{A} = 0$ ,  $\mathbf{A}$  is a  $C^1$ -function, it was shown ([17], based on [13]) that  $(\mathbf{p}_A)^2$  is essentially self-adjoint on  $C_0^\infty(\mathbb{R}^3) \otimes \mathbb{C}^2$ . Later,  $\mathbf{A} \in L_{2,\text{loc}}(\mathbb{R}^3)$  was established as the weakest possible condition for this property to be true [1], [5, p 9]. As a second condition, we require therefore that  $\mathbf{A} \in L_{2,\text{loc}}(\mathbb{R}^3)$ . Let the magnetic field satisfy

$$N_B(\mathbf{x}) := \int_{|\mathbf{x}-\mathbf{y}| \leq 1} |\mathbf{B}(\mathbf{y})|^2 d\mathbf{y} \leq C \quad (1.5)$$

with a constant  $C \in \mathbb{R}$  independent of  $\mathbf{x}$  ((1.5) holds for any  $\mathbf{B} \in L_2(\mathbb{R}^3)$ ). This guarantees the essential self-adjointness of the Pauli operator. The proof is based on the work of Udim [32, theorem 4.2], showing that a consequence of (1.5) is the  $(\mathbf{p}_A)^2$ -boundedness of  $e\sigma\mathbf{B}$  with bound zero. This property establishes the required essential self-adjointness according to the Kato–Rellich theorem [28, theorem X.12].

From the symmetry of  $\sigma\mathbf{p}_A$ , we have  $(\psi, (\sigma\mathbf{p}_A)^2\psi) = \|\sigma\mathbf{p}_A\psi\|^2 \geq 0$  for  $\psi \in C_0^\infty(\mathbb{R}^3) \otimes \mathbb{C}^2$ . Thus,  $(\sigma\mathbf{p}_A)^2$  is a non-negative, self-adjoint operator (by means of closure). It follows [18, theorem 3.35, p 281] that this is also true for

$$E_A := |D_A| = \sqrt{(\sigma\mathbf{p}_A)^2 + m^2} \geq m \quad (1.6)$$

which is the kinetic energy term of the pseudorelativistic operator that will be introduced in section 2.

Due to the positron degrees of freedom, the Dirac operator  $H$  has a spectrum which is unbounded from below. However, in the spectroscopy of static or slowly moving ions, pair creation plays no role. One of the current techniques, used in the field-free case ( $\mathbf{A} = 0$ ), to construct from  $H$  an operator which solely describes the electronic states is the application of a unitary transformation scheme to  $H$  (see, e.g., [7, 15, 30]). A perturbative expansion in the central field strength  $\gamma$  provides pseudorelativistic operators which are block diagonal in the free (i.e.,  $Z = 0$ ) electronic positive and negative spectral subspaces up to a given order in  $\gamma$ . The zero- plus first-order term in this series, the Brown–Ravenhall operator, has obtained widespread interest because it is simply the restriction of  $H$  to the positive spectral subspace. The terms up to second order, comprising the Jansen–Hess operator, provide, however, a much better representation of the bound-state energies [35]. This operator has been proven to be positive with essential spectrum  $\sigma_{\text{ess}} = [m, \infty)$  for sufficiently small  $\gamma$  [4, 12, 14].

If  $\mathbf{A} \neq 0$ , investigations are scarce. It is known that in the absence of the Coulomb field  $V$ , the Dirac operator can be block diagonalized by means of a Foldy–Wouthuysen transformation  $U_0$  [6, section 3.1],

$$U_0 D_A U_0^{-1} = \beta E_A$$

$$U_0 := \left( \frac{m + E_A}{2E_A} \right)^{\frac{1}{2}} + \frac{\beta \alpha \mathbf{p}_A}{(2E_A(m + E_A))^{\frac{1}{2}}}. \quad (1.7)$$

$U_0^{-1}$  is obtained from  $U_0$  by replacing  $\beta \alpha \mathbf{p}_A$  by  $\alpha \mathbf{p}_A \beta = -\beta \alpha \mathbf{p}_A$ . For later use, we note that  $E_A$  commutes with  $U_0$ ,  $[E_A, U_0] = 0$ , because  $[\beta \alpha \mathbf{p}_A, E_A] = [\beta, E_A] \alpha \mathbf{p}_A + \beta [\alpha \mathbf{p}_A, E_A]$  vanishes (the first commutator being zero since  $E_A$  is block diagonal). There are also

a few studies of the ‘magnetic’ Brown–Ravenhall operator showing that this operator is either unbounded from below (if  $\mathbf{A}$  is disregarded in the projector onto the positive spectral subspace [10]) or that it is positive for  $\gamma < \frac{2}{\pi}$  (if  $\mathbf{A}$  is not disregarded) which assures stability of relativistic matter in this model [22, 23].

The aim of the present work is to derive the ‘magnetic’ Jansen–Hess operator  $H^{(2)}$  from the corresponding transformation scheme (section 2), to show under which conditions it is positive (theorem 1, section 4) and to provide criteria for  $\sigma_{\text{ess}} = [m, \infty) + E_f$  to hold (theorem 3, section 6). An auxiliary step is the invariance of the essential spectrum upon removal of the Jansen–Hess potential (theorem 2, section 5). Consequently, theorem 3 also holds for the ‘magnetic’ Brown–Ravenhall operator (which results from dropping the second-order term in  $\gamma$ ). The basic difference from the  $\mathbf{A} = 0$  case in constructing and analysing  $H^{(2)}$  is due to the fact that the kinetic energy operator  $E_A$  is no longer a multiplier in momentum space (as is  $E_{A=0} =: E_p = \sqrt{p^2 + m^2}$ ). Hence, formal techniques have to replace Fourier analysis (sections 2 and 3). Moreover, in contrast to the ‘magnetic’ Brown–Ravenhall operator, the required bounds on  $\gamma$  for self-adjointness and positivity depend nontrivially on the magnetic field. Therefore, these bounds are inferior to the  $\mathbf{A} = 0$  case. With  $\gamma \rightarrow 0$  for  $B \rightarrow \infty$ , our analysis makes the Jansen–Hess operator an unlikely candidate for stability of matter. However, for laboratory magnetic fields up to  $10^{12}$  G this operator should be superior to the ‘magnetic’ Brown–Ravenhall operator regarding electron spectroscopy.

## 2. The transformed Dirac operator

Let us define the projector onto the positive magnetic spectral subspace of the electron (defined by switching off  $V$  but fully including  $\mathbf{A}$ ),

$$\Lambda_{A,+} := \frac{1}{2} \left( 1 + \frac{D_A}{|D_A|} \right). \quad (2.1)$$

For any  $\varphi_+ \in \mathcal{H}_{+,1} := \Lambda_{A,+}(H_1(\mathbb{R}^3) \otimes \mathbb{C}^4)$  (where the Sobolev space  $H_1(\mathbb{R}^3) \otimes \mathbb{C}^4$  is the domain of  $D_A$ ), we have trivially  $\Lambda_{A,+}\varphi_+ = \varphi_+$  and  $D_A\varphi_+ = E_A\varphi_+$ , and one easily verifies that with  $\psi := \begin{pmatrix} u \\ 0 \end{pmatrix}$ ,  $u \in H_1(\mathbb{R}^3) \otimes \mathbb{C}^2$ ,  $\varphi_+$  can be expressed as

$$\varphi_+ = U_0^{-1} \psi \quad (2.2)$$

(namely using (1.7),  $D_A(U_0^{-1}\psi) = U_0^{-1}\beta E_A\psi = U_0^{-1}E_A(\beta\psi) = U_0^{-1}E_A\psi = E_A U_0^{-1}\psi$ ).

Let  $H_V := D_A + V$ . We construct a unitary transformation  $U$  such that the transformed Dirac operator decouples the magnetic spectral subspaces of the electron,

$$U^{-1}H_V U = \Lambda_{A,+}(U^{-1}H_V U)\Lambda_{A,+} + \Lambda_{A,-}(U^{-1}H_V U)\Lambda_{A,-} + E_f, \quad (2.3)$$

with  $\Lambda_{A,+}$  from (2.1) and  $\Lambda_{A,-} = 1 - \Lambda_{A,+}$ . The choice of the projector  $\Lambda_{A,+}$  in (2.3) preserves the gauge invariance of the transformed operator [22]. The field energy  $E_f$  is a constant which is not affected by  $U$ . If one defines  $P_+$  as the projector onto the positive spectral subspace of the Dirac operator  $H_V$ , then (2.3) is equivalent to the condition

$$U^{-1}P_+U = \Lambda_{A,+}. \quad (2.4)$$

If, in addition, the Foldy–Wouthuysen transformation  $U_0$  is applied, the desired block-diagonal operator is obtained as a consequence of  $U_0\Lambda_{A,+}U_0^{-1} = \frac{1}{2}(1 + \beta)$  (see (1.7) and the discussion below):

$$\begin{aligned} M &= \frac{1}{2}(1 + \beta)M\frac{1}{2}(1 + \beta) + \frac{1}{2}(1 - \beta)M\frac{1}{2}(1 - \beta) =: \begin{pmatrix} h & 0 \\ 0 & g \end{pmatrix} \\ M &:= U_0 U^{-1} H_V U U_0^{-1} \end{aligned} \quad (2.5)$$

where  $h, g$  are matrices in  $\mathbb{C}^{2,2}$ .

Rather than solving (2.4) for  $U$  (which was recently achieved in the field-free case [29, 30]), we start from (2.3) and apply a technique [15] which is equivalent to the Douglas–Kroll transformation scheme [7, 16]. We formally expand  $U = \exp(i \sum_{k=1}^{\infty} B_k)$ , where  $B_k$  is an operator which contains the potential  $V$  to  $k$ th order, and we are interested in the transformed operator which is block diagonal up to second order in the potential strength  $\gamma$ . Denoting by  $H^{(2)}$  the second-order solution of (2.3) restricted to  $\mathcal{H}_{+,1}$  (the ‘magnetic’ Jansen–Hess operator) we have, in analogy to the  $\mathbf{A} = 0$  case,

$$H^{(2)} := \Lambda_{A,+} \left\{ D_A + V + \frac{i}{2} [W_1, B_1] + E_f \right\} \Lambda_{A,+} \quad (2.6)$$

with  $W_1 := \Lambda_{A,+} V \Lambda_{A,-} + \Lambda_{A,-} V \Lambda_{A,+}$  being the off-diagonal part of  $V$ .  $B_1$  is determined from the condition

$$W_1 = -i[D_A, B_1]. \quad (2.7)$$

Alternatively, we can obtain  $B_1$  from (2.4). Using the integral representation of  $P_+$  [18, chapter II.1.4] and expanding  $P_+$  in terms of  $V$  by means of the second resolvent identity, we have

$$\begin{aligned} P_+ &= \frac{1}{2} + \frac{1}{2\pi} \int_{-\infty}^{\infty} d\eta \frac{1}{D_A + V + i\eta} \\ &= \Lambda_{A,+} - \frac{1}{2\pi} \int_{-\infty}^{\infty} d\eta \frac{1}{D_A + i\eta} V \frac{1}{D_A + V + i\eta} = \Lambda_{A,+} + F_A + R \end{aligned} \quad (2.8)$$

where  $\Lambda_{A,+}$  and  $F_A$  are the zero- and first-order terms, respectively, while the remainder  $R$  is of higher order in  $V$ . Defining  $\tilde{D}_A := D_A/|D_A|$  and solving (2.4) up to first order in  $V$ , we get

$$2F_A - iB_1\tilde{D}_A + i\tilde{D}_AB_1 = 0. \quad (2.9)$$

Multiplication of (2.9) by  $\tilde{D}_A$  from the left and, respectively, from the right and addition of the resulting equations provides the useful relation

$$F_A\tilde{D}_A = -\tilde{D}_AF_A. \quad (2.10)$$

Whereas (2.9) is also only an implicit equation for  $B_1$ , a trial for  $B_1$  can be found from the formal solution  $U$  of (2.4) which is completely analogous to the field-free case [29],  $U^{-1} = [1 + (\Lambda_{A,+} - \Lambda_{A,-})(P_+ - \Lambda_{A,+})](1 - (P_+ - \Lambda_{A,+})^2)^{-\frac{1}{2}}$ . An expansion of this formal solution up to first order in  $V$  leads to

$$B_1 = i\tilde{D}_AF_A. \quad (2.11)$$

With the help of (2.10), it is easily verified that (2.11) is indeed a solution to (2.9). Insertion into (2.6) finally results in

$$\begin{aligned} H^{(2)} &= \Lambda_{A,+} \{ D_A + V + B_{2m} + E_f \} \Lambda_{A,+} \\ B_{2m} &:= \frac{1}{4} [V F_A \tilde{D}_A + \tilde{D}_A F_A V + \tilde{D}_A V F_A + F_A V \tilde{D}_A]. \end{aligned} \quad (2.12)$$

### 3. Relative form boundedness of the Jansen–Hess potential

In order to establish self-adjointness of  $H^{(2)}$ , the form boundedness of the potential contributions to  $H^{(2)}$  (restricted to the ‘positive’ space  $\mathcal{H}_{+,1}$ ) relative to the kinetic energy operator  $E_A$  is needed. We have to fix the potential strength  $\gamma$  such that this bound becomes smaller than one. We start by showing the relative boundedness of the linear term (in  $\gamma$ )  $V$ , then we prove the boundedness of the operator  $B_1$  (introduced by the transformation  $U$ ) and subsequently the relative boundedness of the quadratic term. The resulting form boundedness

of the Jansen–Hess potential relative to  $E_A$  is stated in lemma 1, and the condition for  $H^{(2)}$  being self-adjoint is part of theorem 1.

3.1.  $E_A$ -boundedness of  $V$  and boundedness of  $B_1$

A basic ingredient is the inequality  $(\varphi, \exp(-\mathbf{p}_A^2 t)\varphi) \leq (\varphi, \exp(-p^2 t)\varphi)$ , valid for  $t \geq 0$  and  $\mathbf{A} \in L_{2,\text{loc}}(\mathbb{R}^3)$  ([1] and references therein). Making use of  $(\varphi, p^2 \varphi) = -\lim_{t \rightarrow 0} (\varphi, \frac{\exp(-tp^2)-1}{t}\varphi)$  [24], one derives

$$(\varphi, (\mathbf{p} - e\mathbf{A})^2 \varphi) \geq (\varphi, p^2 \varphi) \tag{3.1}$$

which is known as diamagnetic inequality (see also earlier work [13] for the related inequality  $(|\varphi|, p^2 |\varphi|) \leq (\varphi, (\mathbf{p} - e\mathbf{A})^2 \varphi)$ ). A consequence is

$$|\mathbf{p} - e\mathbf{A}| \geq p. \tag{3.2}$$

Further, let  $\mathcal{O}_- := \frac{1}{2}(|\mathcal{O}| - \mathcal{O}) \geq 0$  be the negative part of an operator  $\mathcal{O}$  and  $\text{tr } \mathcal{O}_-$  its trace (i.e., the sum over the absolute values of the negative eigenvalues of  $\mathcal{O}$  times the spin degrees of freedom). Then by means of (3.1) and the Lieb–Thirring inequality [21, 23] for any  $\mu > 0$  and  $d > 0$  one has

$$\text{tr}[\mu(\mathbf{p} - e\mathbf{A})^2 + e\sigma\mathbf{B}]_-^d \leq \mu^d \text{tr} \left[ p^2 + \frac{e\sigma\mathbf{B}}{\mu} \right]_-^d \leq 2\mu^d L_{d,3} \int_{\mathbb{R}^3} \left( \frac{e|\mathbf{B}|}{\mu} \right)^{d+\frac{3}{2}} \text{d}\mathbf{x} \tag{3.3}$$

with constants  $L_{\frac{1}{2},3} \leq 0.06003$  and  $L_{1,3} \leq 0.0403$ .

Then, following [23] we get the form estimate for  $\varphi \in H_1(\mathbb{R}^3) \otimes \mathbb{C}^4$ ,  $\|\varphi\| = 1$ , using Kato’s inequality  $\frac{1}{x} \leq \frac{\pi}{2} p$  and (3.2) as well as the trace inequality for non-negative, self-adjoint operators,  $\text{tr}(\mathcal{O}_1 - \mathcal{O}_2)_- \leq \text{tr}(\mathcal{O}_1^2 - \mathcal{O}_2^2)_-^{\frac{1}{2}}$ ,

$$\begin{aligned} (\varphi, E_A \varphi) - \left( \varphi, \frac{\gamma_0}{x} \varphi \right) &\geq (\varphi, \sqrt{E_A^2 - m^2} \varphi) - \frac{\gamma_0 \pi}{2} (\varphi, |\mathbf{p} - e\mathbf{A}| \varphi) \\ &\geq -\text{tr} \left[ (E_A^2 - m^2) - \left( \frac{\gamma_0 \pi}{2} |\mathbf{p} - e\mathbf{A}| \right)^2 \right]_-^{\frac{1}{2}} \\ &\geq -2L_{\frac{1}{2},3} \frac{e^2}{[1 - (\gamma_0 \pi / 2)^2]^{\frac{3}{2}}} \|\mathbf{B}\|^2 \end{aligned} \tag{3.4}$$

for  $\gamma_0 < \frac{2}{\pi}$ .

Moreover, using  $\text{tr } \mathcal{O}_- \leq (\text{tr } \mathcal{O}_-^{\frac{1}{2}})^2$  [27, p 210] and Hardy’s inequality  $\frac{1}{x^2} \leq 4p^2$ ,

$$\begin{aligned} \|E_A \varphi\|^2 - \left\| \frac{\gamma_1}{x} \varphi \right\|^2 &\geq (\varphi, [(1 - 4\gamma_1^2)(\mathbf{p} - e\mathbf{A})^2 - e\sigma\mathbf{B}] \varphi) \\ &\geq -(\text{tr} [(1 - 4\gamma_1^2)(\mathbf{p} - e\mathbf{A})^2 - e\sigma\mathbf{B}]_-^{\frac{1}{2}})^2 \\ &\geq - \left[ 2L_{\frac{1}{2},3} \frac{e^2}{[1 - 4\gamma_1^2]^{\frac{3}{2}}} \|\mathbf{B}\|^2 \right]^2 \end{aligned} \tag{3.5}$$

for  $\gamma_1 < \frac{1}{2}$ . Thus, we obtain the  $E_A$ -boundedness of the potential  $V$  in the form and in the norm [28, p 162],

$$|(\varphi, V \varphi)| \leq \frac{\gamma}{\gamma_0} (\varphi, E_A \varphi) + \gamma c_B (\varphi, \varphi), \quad c_B := \frac{2}{\gamma_0} L_{\frac{1}{2},3} \frac{e^2}{[1 - (\gamma_0 \pi / 2)^2]^{\frac{3}{2}}} \|\mathbf{B}\|^2 \tag{3.6}$$

and

$$\|V\varphi\| \leq \frac{\gamma}{\gamma_1} \|E_A\varphi\| + \gamma d_B \|\varphi\|, \quad d_B := \frac{2}{\gamma_1} L_{\frac{1}{2},3} \frac{e^2}{[1 - 4\gamma_1^2]^{\frac{3}{2}}} \|\mathbf{B}\|^2. \quad (3.7)$$

The boundedness of  $B_1$  is a consequence of the boundedness of  $F_A$ , since

$$\|B_1\| \leq \|\tilde{D}_A\| \|F_A\| = \|F_A\|. \quad (3.8)$$

With (3.6) at hand, the boundedness of  $F_A$  is easy to show. Following the proof of [30, lemma 1], we have for  $\varphi_+, \psi_+ \in \mathcal{H}_{+,1}$  from (2.8)

$$\begin{aligned} \|F_A\| &= \frac{1}{2\pi} \left\| \int_{-\infty}^{\infty} d\eta \frac{1}{D_A + i\eta} V \frac{1}{D_A + i\eta} \right\| \\ &\leq \frac{\gamma}{2\pi} \sup_{\|\varphi_+\|=\|\psi_+\|=1} \int_{-\infty}^{\infty} d\eta \left| \left( \varphi_+, \frac{1}{D_A + i\eta} \frac{1}{x^{1/2}} \cdot \frac{1}{x^{1/2}} \frac{1}{D_A + i\eta} \psi_+ \right) \right| \\ &\leq \frac{\gamma}{2\pi} \sup_{\|\varphi_+\|=\|\psi_+\|=1} \int_{-\infty}^{\infty} d\eta \left\| \frac{1}{x^{1/2}} \frac{1}{D_A - i\eta} \varphi_+ \right\| \cdot \left\| \frac{1}{x^{1/2}} \frac{1}{D_A + i\eta} \psi_+ \right\|. \end{aligned} \quad (3.9)$$

An application of the Schwarz inequality leads to

$$\|F_A\| \leq \frac{\gamma}{2\pi} \sup_{\|\varphi_+\|=\|\psi_+\|=1} \left( \int_{-\infty}^{\infty} d\eta \left\| \frac{1}{x^{1/2}} \frac{1}{D_A - i\eta} \varphi_+ \right\|^2 \right)^{\frac{1}{2}} \cdot \left( \int_{-\infty}^{\infty} d\eta \left\| \frac{1}{x^{1/2}} \frac{1}{D_A + i\eta} \psi_+ \right\|^2 \right)^{\frac{1}{2}}. \quad (3.10)$$

Setting  $\varphi := \frac{1}{D_A - i\eta} \varphi_+$  (note that  $D_A^2 > 0$  for  $m \neq 0$  such that  $(D_A - i\eta)^{-1}$  is bounded for  $\eta \in \mathbb{R}$ ), we have from (3.6)

$$\left\| \frac{1}{x^{1/2}} \frac{1}{D_A - i\eta} \varphi_+ \right\|^2 = \left( \varphi, \frac{1}{x} \varphi \right) \leq \frac{1}{\gamma_0} (\varphi, E_A \varphi) + c_B (\varphi, \varphi) \quad (3.11)$$

and thus we get for the two (equal) integrals in (3.10), using  $D_A \varphi_+ = E_A \varphi_+$ ,

$$\begin{aligned} \int_{-\infty}^{\infty} d\eta \left\| \frac{1}{x^{1/2}} \frac{1}{D_A - i\eta} \varphi_+ \right\|^2 &\leq \frac{1}{\gamma_0} \left( \varphi_+, \int_{-\infty}^{\infty} d\eta \frac{1}{E_A + i\eta} E_A \frac{1}{E_A - i\eta} \varphi_+ \right) \\ &\quad + c_B \left( \varphi_+, \int_{-\infty}^{\infty} d\eta \frac{1}{E_A^2 + \eta^2} \varphi_+ \right) = \frac{1}{\gamma_0} \cdot \pi \|\varphi_+\|^2 + c_B \pi \left( \varphi_+, \frac{1}{E_A} \varphi_+ \right). \end{aligned} \quad (3.12)$$

We estimate  $E_A \geq m$  and finally obtain the boundedness of  $\|F_A\|$ :

$$\|F_A\| \leq \frac{\gamma}{2\gamma_0} \left( 1 + c_B \frac{\gamma_0}{m} \right). \quad (3.13)$$

We note that due to the existence of zero modes [25] the lower bound  $m$  of  $E_A$  is sharp: there is a field  $\mathbf{B}_0 = \nabla \times \mathbf{A}_0 \in L_2(\mathbb{R}^3)$ , satisfying (1.4) and hence (1.5), and a function  $\psi_0 \in H_1(\mathbb{R}^3) \setminus \{0\} \otimes \mathbb{C}^2$  such that

$$\sigma(\mathbf{p} - e\mathbf{A}_0)\psi_0 = 0. \quad (3.14)$$

From this it follows that the 4-spinor  $\begin{pmatrix} \psi_0 \\ 0 \end{pmatrix}$  obeys  $D_{A_0} \begin{pmatrix} \psi_0 \\ 0 \end{pmatrix} = m \begin{pmatrix} \psi_0 \\ 0 \end{pmatrix}$ , i.e. it lies in the positive magnetic spectral subspace of the electron, and  $m$  is the lowest positive eigenvalue of  $D_{A_0}$ .

3.2. Relative boundedness of the Jansen–Hess potential

From (2.12) we get for  $\psi_+ \in \mathcal{H}_{+,1}$ , with  $\|\Lambda_{A,+}\| = 1$ ,

$$\begin{aligned} \|(4\Lambda_{A,+}B_{2m}\Lambda_{A,+})\psi_+\| &\leq \|4B_{2m}\psi_+\| \\ &\leq \|VF_A\tilde{D}_A\psi_+\| + \|\tilde{D}_AF_AV\psi_+\| + \|\tilde{D}_AVF_A\psi_+\| + \|F_AV\tilde{D}_A\psi_+\|. \end{aligned} \tag{3.15}$$

We shall estimate each of these four terms separately, using the boundedness (3.13) of  $F_A$  and the relative boundedness (3.7) of  $V$ . First, we show

$$[D_A, F_A] = \frac{1}{2}[\tilde{D}_A, V]. \tag{3.16}$$

We multiply (2.7) with  $\tilde{D}_A$  and insert  $B_1$  from (2.11). This gives

$$\tilde{D}_AW_1 = -i\tilde{D}_A(iD_A\tilde{D}_AF_A - i\tilde{D}_AF_AD_A) = [D_A, F_A]. \tag{3.17}$$

Inserting for  $W_1$  (below (2.6)) results in (3.16).

Using that  $\|\tilde{D}_A\| = 1$  and  $\tilde{D}_A\psi_+ = \psi_+$ , (3.15) gives

$$\|4B_{2m}\psi_+\| \leq 2\|VF_A\psi_+\| + 2\|F_A\|\|V\psi_+\|. \tag{3.18}$$

With (3.7) and (3.16), defining  $F_A\psi_+ =: \varphi$ , we estimate the first term by

$$\begin{aligned} \|VF_A\psi_+\| &\leq \frac{\gamma}{\gamma_1}\|D_A\varphi\| + \gamma d_B\|\varphi\| \leq \frac{\gamma}{\gamma_1}\|D_AF_A\psi_+\| + \gamma d_B\|F_A\|\|\psi_+\| \\ &\leq \frac{\gamma}{\gamma_1}\left\{\|F_A\|\|D_A\psi_+\| + \frac{1}{2}\|\tilde{D}_A\|\|V\psi_+\| + \frac{1}{2}\|V\psi_+\|\right\} + \gamma d_B\|F_A\|\|\psi_+\|. \end{aligned} \tag{3.19}$$

Thus, we get

$$\|B_{2m}\psi_+\| \leq \frac{\gamma}{\gamma_1}\left(\frac{\gamma}{2\gamma_1} + \|F_A\|\right)\|D_A\psi_+\| + \gamma d_B\left(\frac{\gamma}{2\gamma_1} + \|F_A\|\right)\|\psi_+\|. \tag{3.20}$$

Using (3.13) this results in

$$\begin{aligned} \|B_{2m}\psi_+\| &\leq c\|E_A\psi_+\| + C\|\psi_+\| \\ c &:= \frac{\gamma^2}{2\gamma_1}\left(\frac{1}{\gamma_1} + \frac{1}{\gamma_0} + \frac{c_B}{m}\right), \quad C := \gamma^2\frac{d_B}{2}\left(\frac{1}{\gamma_1} + \frac{1}{\gamma_0} + \frac{c_B}{m}\right). \end{aligned} \tag{3.21}$$

Note that both constants,  $c$  and  $C$ , depend on the field energy through  $\|\mathbf{B}\| = (8\pi E_f)^{\frac{1}{2}}$ .

From the  $E_A$ -boundedness of  $B_{2m}$  follows the  $E_A$ -form boundedness of  $B_{2m}$  with the same relative bound  $c$  [28, p 168]. Thus, we have proven

**Lemma 1.** *Let  $H^{(2)} = D_A + V + B_{2m} + E_f$  be the ‘magnetic’ Jansen–Hess operator acting on  $\mathcal{H}_{+,1}$ . Then  $V + B_{2m}$  is  $E_A$ -form bounded,*

$$|(\psi_+, (V + B_{2m})\psi_+)| \leq \left(\frac{\gamma}{\gamma_0} + c\right)(\psi_+, E_A\psi_+) + \tilde{C}(\psi_+, \psi_+), \tag{3.22}$$

with  $c = \frac{\gamma^2}{2\gamma_1}\left(\frac{1}{\gamma_1} + \frac{1}{\gamma_0} + \frac{c_B}{m}\right)$ , where  $c_B$  and  $d_B$  are defined in (3.6) and (3.7), and  $\tilde{C}$  is some  $\|\mathbf{B}\|$ -dependent constant. (The parameters  $\gamma_0 < \frac{2}{\pi}$  and  $\gamma_1 < \frac{1}{2}$  can be chosen arbitrarily.)



#### 4. Positivity of $H^{(2)}$

Let  $\delta > 0$  and recall that  $E_A \geq m$  is bounded below. If in (3.21), the  $\delta E_A$ -bound  $\frac{c}{\delta}$  of  $B_{2m}$  is smaller than unity, then according to [18, theorem 4.11, p 291]  $\delta E_A + B_{2m}$  is also bounded below by means of

$$(\psi_+, (\delta E_A + B_{2m})\psi_+) \geq \left( \delta m - \max \left\{ \frac{C}{1 - c/\delta}, C + cm \right\} \right) (\psi_+, \psi_+), \quad (4.1)$$

where the constants  $c$  and  $C$  are defined in (3.21).

Using the above results, we can estimate

$$\begin{aligned} (\psi_+, H^{(2)}\psi_+) &= (\psi_+, E_A\psi_+) - |(\psi_+, V\psi_+)| + (\psi_+, B_{2m}\psi_+) + E_f(\psi_+, \psi_+) \\ &\geq \left( \psi_+, \left( \left( 1 - \frac{\gamma}{\gamma_0} \right) E_A + B_{2m} \right) \psi_+ \right) - \gamma c_B(\psi_+, \psi_+) + E_f(\psi_+, \psi_+) \\ &\geq \left( \left( 1 - \frac{\gamma}{\gamma_0} \right) m - \max \left\{ \frac{C}{1 - c/(1 - \gamma/\gamma_0)}, C + cm \right\} - \gamma c_B + E_f \right) (\psi_+, \psi_+). \end{aligned} \quad (4.2)$$

This results in

**Theorem 1.** *Let  $H^{(2)} = D_A + V + B_{2m} + E_f$  be the ‘magnetic’ Jansen–Hess operator acting on  $\mathcal{H}_{+,1}$ . If the  $E_A$ -form bound of  $V + B_{2m}$  is smaller than unity,*

$$\frac{\gamma}{\gamma_0} + c < 1, \quad (4.3)$$

*then  $H^{(2)}$  is bounded below and thus extends to a self-adjoint operator on  $\Lambda_{A,+}(L_2(\mathbb{R}^3) \otimes \mathbb{C}^4)$ . If in addition*

$$\left( 1 - \frac{\gamma}{\gamma_0} \right) m - \gamma c_B - \max \left\{ \frac{C(1 - \gamma/\gamma_0)}{1 - \gamma/\gamma_0 - c}, C + cm \right\} + E_f > 0, \quad (4.4)$$

*then  $H^{(2)}$  is positive. This restricts the potential strength to  $\gamma < \gamma_c$  where  $\gamma_c \leq 0.353$  depending on the magnetic field  $\mathbf{B}$ .*

In order to derive the conditions on the bound for  $\gamma$  which are required for theorem 1, we first consider the case  $\mathbf{B} = 0$ . Then, we can set  $\gamma_0 = \frac{2}{\pi}$  and  $\gamma_1 = \frac{1}{2}$ , and both inequalities, (4.3) and (4.4), are satisfied for  $\gamma < \gamma_c^{(0)}$ , where  $\gamma_c^{(0)} = 0.353 (Z \leq 48)$  is a solution of

$$\frac{\gamma}{\gamma_0} + c = \gamma \frac{\pi}{2} + \gamma^2 \left( 2 + \frac{\pi}{2} \right) = 1. \quad (4.5)$$

This is considerably smaller than the critical  $\gamma$  obtained earlier for the field-free case ( $\gamma_c^{(0)} = 1.006$  [4]), where one is able to work in momentum space and to use Mellin transform techniques.

When  $\mathbf{B}$  is turned on, the bound on  $\gamma$  from the self-adjointness condition decreases slowly. For example, if  $\|\mathbf{B}\| = 2.5$ , then by optimizing  $\gamma_0$  and  $\gamma_1$  one gets from (4.3)  $\gamma_c = 0.335 (\gamma_0 = 0.6, \gamma_1 = 0.498)$ , whereas positivity is guaranteed for  $\gamma < 0.316 (\gamma_0 = 0.6, \gamma_1 = 0.47)$ . The relativistic ground-state binding energy of an electron,  $|E_g - m| := m|\sqrt{1 - \gamma^2} - 1| = 0.0644$  (in units where  $m = 1$ , using  $\gamma = \gamma_c^{(0)}$ ), may be used as a reference value with which to compare the field energy  $E_f$ . Even for quite large fields<sup>1</sup>, e.g.  $\|\mathbf{B}\| = 10$  (where  $E_f \approx 60|E_g - m|$ ), the critical potential strength (with

<sup>1</sup> In conventional units,  $B = 1m^2e^3c/\hbar^3 = 2.35 \times 10^9$  G.

$\gamma_0 = 0.54, \gamma_1 = 0.499$ ) has only slightly decreased,  $\gamma_c = 0.299$  ( $Z < 41$ ) while  $H^{(2)} > 0$  for  $\gamma < 0.275$  ( $\gamma_0 = 0.54, \gamma_1 = 0.45$ ).

However, when  $\|\mathbf{B}\|$  becomes extremely large (but still is finite), our estimates (resulting in (4.4)) no longer guarantee positivity because  $C$  is of fourth order in  $\|\mathbf{B}\|$  and eventually dominates  $E_f$ . In order to remedy this deficiency, different estimates for the  $E_A$ -boundedness of the potential  $V$  are required.

For the magnetic fields which are  $(\mathbf{p}_A)^2$ -bounded with bound  $\kappa \rightarrow 0$  (and hence also  $(\mathbf{p}_A)^2$ -form bounded with the same bound), we have from (1.3)

$$\begin{aligned} (\varphi, |\mathbf{B}|\varphi) &\leq \kappa(\varphi, \mathbf{p}_A^2 \varphi) + C_\kappa(\varphi, \varphi) \\ &\leq \kappa(\varphi, E_A^2 \varphi) + \kappa e(\varphi, |\mathbf{B}|\varphi) + C_\kappa(\varphi, \varphi) \end{aligned} \tag{4.6}$$

proving the  $E_A^2$ -form boundedness of  $|\mathbf{B}|$  with bound  $\kappa/(1 - \kappa e)$ . It can be shown [34, proof of theorem 10.17] that the constant  $C_\kappa$  depends linearly on  $\sup_{\mathbf{x} \in \mathbb{R}^3} (N_B(\mathbf{x}))^{\frac{1}{2}}$  which in turn can be estimated above by  $\|\mathbf{B}\|$ . So, we get from Hardy’s inequality and (3.1)

$$(\varphi, V^2 \varphi) \leq 4\gamma^2(\varphi, \mathbf{p}_A^2 \varphi) \leq 4\gamma^2(\varphi, E_A^2 \varphi) + 4\gamma^2 e(\varphi, |\mathbf{B}|\varphi). \tag{4.7}$$

Using (4.6), we eventually obtain the estimate

$$\|V\varphi\| \leq 2\gamma \left(1 + \frac{\kappa e}{1 - \kappa e}\right)^{\frac{1}{2}} \|E_A \varphi\| + 2\gamma \left(\frac{e}{1 - \kappa e}\right)^{\frac{1}{2}} C_\kappa^{\frac{1}{2}} \|\varphi\| \tag{4.8}$$

in place of (3.7). Note that  $\kappa$  can be taken arbitrarily close to 0 such that the  $E_A$ -bound of  $V$  agrees with the one in (3.7). However, the last term in (4.8) increases only  $\sim \|\mathbf{B}\|^{\frac{1}{2}}$ . A similar estimate replaces (3.6) for the  $E_A$ -form boundedness of  $V$ .

In order to get explicit constants, let us for the moment assume that  $\mathbf{B}$  is bounded with  $\|\mathbf{B}\|_\infty \leq \|\mathbf{B}\|$ . Then, the last term in (4.7) is estimated by  $4\gamma^2 e \|\mathbf{B}\|_\infty(\varphi, \varphi) \leq 4\gamma^2 e \|\mathbf{B}\|(\varphi, \varphi)$ , giving  $\|V\varphi\| \leq 2\gamma \|E_A \varphi\| + 2\gamma (e \|\mathbf{B}\|)^{\frac{1}{2}} \|\varphi\|$ . For the form bound, using Kato’s inequality, one gets

$$\begin{aligned} |(\varphi, V\varphi)| &\leq \gamma \frac{\pi}{2}(\varphi, |\mathbf{p} - e\mathbf{A}|\varphi) \leq \gamma \frac{\pi}{2}(\varphi, \sqrt{E_A^2 + e|\mathbf{B}|}\varphi) \\ &\leq \gamma \frac{\pi}{2}(\varphi, E_A \varphi) + \gamma \frac{\pi}{2}(e \|\mathbf{B}\|)^{\frac{1}{2}}(\varphi, \varphi). \end{aligned} \tag{4.9}$$

When (3.6) and (3.7) are replaced by these two inequalities in the subsequent estimates, conditions (4.3) and (4.4) of theorem 1 now read

$$1 - \gamma \frac{\pi}{2} - c_1 > 0 \tag{4.10}$$

and

$$\left(1 - \gamma \frac{\pi}{2}\right) m - \gamma \frac{\pi}{2}(e \|\mathbf{B}\|)^{\frac{1}{2}} - \max \left\{ \frac{C_1(1 - \gamma\pi/2)}{1 - \gamma\pi/2 - c_1}, C_1 + c_1 m \right\} + E_f > 0 \tag{4.11}$$

where  $c_1$  and  $C_1$  are the changed bounds for  $B_{2m}$ , replacing (3.21),

$$c_1 := \gamma^2 \left(2 + \frac{\pi}{2} + \frac{\pi}{2} \frac{(e \|\mathbf{B}\|)^{\frac{1}{2}}}{m}\right), \quad C_1 := \gamma^2 \left(\left[2 + \frac{\pi}{2}\right] (e \|\mathbf{B}\|)^{\frac{1}{2}} + \frac{\pi}{2} \frac{e \|\mathbf{B}\|}{m}\right). \tag{4.12}$$

In condition (4.11) for the positivity of  $H^{(2)}$  the leading term in  $\|\mathbf{B}\|$  is now  $E_f$ , guaranteeing positivity for sufficiently large  $\|\mathbf{B}\|$ . For example, for  $\|\mathbf{B}\| = 10$ , (4.10) and (4.11) hold for  $\gamma < 0.304$ , this limit already exceeding the corresponding one from (4.3).

We close this section by showing that a  $\mathbf{B}$ -dependent constant in the form boundedness of  $V$  (which in turn leads to a  $\mathbf{B}$ -dependent condition (4.3) for self-adjointness of  $H^{(2)}$ ) cannot be avoided [36].

It was proven [2] that for a homogeneous magnetic field  $\mathbf{B}$ , the ground-state energy of the Pauli operator in a central Coulomb field of any given strength  $Z_0 e^2$  diverges logarithmically with  $B$ . This leads to the estimate

$$\frac{1}{2m}(\varphi, (E_A^2 - m^2)\varphi) - \left(\varphi, \frac{Z_0 e^2}{x}\varphi\right) \geq -c_0(\ln B)^2(\varphi, \varphi), \quad (4.13)$$

with a suitable ( $Z_0$ -dependent) constant  $c_0$  and sufficiently large  $B$ . The estimate is sharp since (4.13) turns into an equality if  $\varphi$  is the ground-state function. Let  $Z_0 = Z/2$ . Then, (4.13) is written in the following way:

$$\begin{aligned} \left(\varphi, \frac{2Z_0 e^2}{x}\varphi\right) &= |(\varphi, V\varphi)| \\ &\leq c_3(\varphi, E_A\varphi) + \left(\varphi, E_A\left(\frac{E_A}{m} - c_3\right)\varphi\right) + [2c_0(\ln B)^2 - m](\varphi, \varphi), \end{aligned} \quad (4.14)$$

where  $0 < c_3 < 1$  is an arbitrary real number. Since  $E_A \geq m$ , the second term in (4.14) is positive and cannot compensate the  $B$ -dependence of the third term for  $B \rightarrow \infty$ . The fact that a homogeneous  $\mathbf{B}$ -field violates our requirement  $\|\mathbf{B}\| < \infty$  is no serious problem, since the strong localization of the ground-state function in all three spatial directions [2, 26] allows for the replacement of the homogeneous  $\mathbf{B}$  by an  $L_2$ -field (by smoothly cutting off at very large distances) without changing the ground-state energy.

## 5. Relative compactness of the perturbation

The aim of this section is to prove

**Theorem 2.** *Let  $H^{(2)} = H_0 + W$  be the ‘magnetic’ Jansen–Hess operator with  $H_0 := \Lambda_{A,+}(D_A + E_f)\Lambda_{A,+}$  and  $W := \Lambda_{A,+}(V + B_{2m})\Lambda_{A,+}$ . Then, we have for  $\gamma < \tilde{\gamma}_c$*

$$\sigma_{\text{ess}}(H^{(2)}) = \sigma_{\text{ess}}(H_0). \quad (5.1)$$

The critical potential strength is  $\tilde{\gamma}_c \leq \tilde{\gamma}_c^{(0)} = 0.319$  and depends on the magnetic field  $\mathbf{B}$ .

Equivalently [18, problem 5.38, p 244], we have to prove the compactness of the difference  $R_\mu$  of the resolvents of  $H^{(2)}$  and  $H_0$ ,

$$R_\mu := \frac{1}{H^{(2)} + \mu} - \frac{1}{H_0 + \mu} = -\frac{1}{H_0 + \mu}\Lambda_{A,+}(V + B_{2m})\Lambda_{A,+}\frac{1}{H^{(2)} + \mu}, \quad (5.2)$$

where the second resolvent identity is used, and  $\mu > 0$  has to be chosen suitably. We decompose

$$\begin{aligned} R_\mu &=: R_\mu(V) + R_\mu(B_{2m}) \\ &= -\left\{\frac{1}{H_0 + \mu}(\Lambda_{A,+}V\Lambda_{A,+} + \Lambda_{A,+}B_{2m}\Lambda_{A,+})\frac{1}{(H_0 + \mu)^\lambda}\right\}\left[(H_0 + \mu)^\lambda\frac{1}{H^{(2)} + \mu}\right] \end{aligned} \quad (5.3)$$

where  $\lambda \in \{\frac{1}{2}, 1\}$ , and we will show that the two operators in curly brackets are compact while the factor in square brackets is bounded. This will prove the compactness of  $R_\mu$ .

### 5.1. Relative compactness of $V^{\frac{1}{2}}$

For the proof of the above assertion we need, with  $V = -\gamma/x$ , the following lemma.

**Lemma 2.** *Let  $H_0 = \Lambda_{A,+}(D_A + E_f)\Lambda_{A,+}$  with  $D_A$  from (1.1) and  $\Lambda_{A,+}$  from (2.1). Then, the operator*

$$\frac{1}{x^{\frac{1}{2}}}\Lambda_{A,+}\frac{1}{H_0 + \mu} \tag{5.4}$$

is compact for  $\mu > 0$ .

According to [18, theorem 4.10, p 159], its adjoint  $(H_0 + \mu)^{-1}\Lambda_{A,+}x^{-\frac{1}{2}}$  is then compact too.

**Proof.** We start by showing the boundedness of  $x^{-\frac{1}{2}}(|D_A| + \mu)^{-\frac{1}{2}}$  on  $L_2(\mathbb{R}^3) \otimes \mathbb{C}^4$ . From (3.6), we get

$$\left\| \frac{1}{x^{\frac{1}{2}}}\frac{1}{(|D_A| + \mu)^{\frac{1}{2}}}\psi \right\|^2 \leq \frac{1}{\gamma_0} \left\| |D_A|^{\frac{1}{2}}\frac{1}{(|D_A| + \mu)^{\frac{1}{2}}}\psi \right\|^2 + c_B \left\| \frac{1}{(|D_A| + \mu)^{\frac{1}{2}}}\psi \right\|^2. \tag{5.5}$$

Since  $(|D_A| + \mu)^{-\frac{1}{2}}$  is bounded for  $\mu > 0$  and since  $|D_A|(|D_A| + \mu)^{-1} \leq 1$ , the rhs of (5.5) is bounded. This implies the relative boundedness of  $x^{-\frac{1}{2}}$  with respect to  $|D_A|$  with form bound  $a = 0$ . In fact, using [28, p 340, problem 19],

$$a = \lim_{\mu \rightarrow \infty} \left\| \frac{1}{x^{\frac{1}{2}}}(|D_A| + \mu)^{-1} \right\|, \tag{5.6}$$

we have from (5.5), with  $|D_A| \geq m$ ,

$$\|x^{-\frac{1}{2}}(|D_A| + \mu)^{-1}\psi\| \leq \|x^{-\frac{1}{2}}(|D_A| + \mu)^{-\frac{1}{2}}\| \left(\frac{1}{m + \mu}\right)^{\frac{1}{2}} \|\psi\|$$

which proves  $a = 0$ .

Following [31, lemma 11.5], we define a smooth function  $\chi_0 \in C_0^\infty(\mathbb{R}^3)$  mapping to  $[0, 1]$  by means of

$$\chi_0(\mathbf{x}) := \begin{cases} 1, & x < R \\ 0, & x \geq R + 1 \end{cases} \tag{5.7}$$

with some  $R > 0$ , such that  $\text{supp}(1 - \chi_0) \subset \mathbb{R}^3 \setminus B_R(0)$ , where  $B_R(0)$  is a ball of radius  $R$  centred at the origin. Further, let  $(\psi_n)_{n \in \mathbb{N}}$  be a normalized sequence in  $H_1(\mathbb{R}^3) \otimes \mathbb{C}^4$  weakly converging to zero. We prove the compactness of (5.4) by showing that  $\|x^{-\frac{1}{2}}\Lambda_{A,+}(H_0 + \mu)^{-1}\psi_n\| \rightarrow 0$  for  $n \rightarrow \infty$ . We decompose

$$\left\| \frac{1}{x^{\frac{1}{2}}}\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\| \leq \left\| (1 - \chi_0)\frac{1}{x^{\frac{1}{2}}}\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\| + \left\| \frac{1}{x^{\frac{1}{2}}}\chi_0\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\|. \tag{5.8}$$

For the first term, we have

$$\left\| (1 - \chi_0)\frac{1}{x^{\frac{1}{2}}}\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\| \leq \frac{1}{R^{\frac{1}{2}}} \left\| (1 - \chi_0)\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\| \leq \frac{c}{R^{\frac{1}{2}}} \tag{5.9}$$

with some constant  $c$ . Thus, it can be made smaller than  $\epsilon/2$  if  $R > (2c/\epsilon)^2$ .

For the second term, we define  $\tilde{\psi}_n := \chi_0\Lambda_{A,+}(H_0 + \mu)^{-1}\psi_n$  and use the  $|D_A|$ -boundedness of  $x^{-\frac{1}{2}}$  with bound  $a \rightarrow 0$ ,

$$\left\| \frac{1}{x^{\frac{1}{2}}}\chi_0\Lambda_{A,+}\frac{1}{H_0 + \mu}\psi_n \right\| \leq a\| |D_A|\tilde{\psi}_n\| + b\|\tilde{\psi}_n\|, \tag{5.10}$$

with some constant  $b$ . In order to establish that  $\| |D_A|\tilde{\psi}_n\|$  is finite (such that  $a\| |D_A|\tilde{\psi}_n\|$  can be dropped), we consider

$$\mathcal{O} := [D_A, \chi_0] = \alpha(\mathbf{p}\chi_0) \tag{5.11}$$

which is bounded because  $\chi_0$  is a  $C_0^\infty$ -function. Thus, we can decompose

$$\chi_0 |D_A|^2 \chi_0 = \chi_0 D_A \cdot D_A \chi_0 = D_A \chi_0^2 D_A - \mathcal{O} \chi_0 D_A + D_A \chi_0 \mathcal{O} - \mathcal{O}^2 \quad (5.12)$$

and estimate

$$\begin{aligned} \| |D_A| \tilde{\psi}_n \|^2 &= \left( \psi_n, \frac{1}{H_0 + \mu} \Lambda_{A,+} (D_A \chi_0^2 D_A - \mathcal{O} \chi_0 D_A + D_A \chi_0 \mathcal{O} - \mathcal{O}^2) \Lambda_{A,+} \frac{1}{H_0 + \mu} \psi_n \right) \\ &\leq \| \chi_0 \|_\infty^2 \left\| D_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \psi_n \right\|^2 + \left\| \mathcal{O} \Lambda_{A,+} \frac{1}{H_0 + \mu} \psi_n \right\|^2 \\ &\quad \times \| \chi_0 \|_\infty \left\| D_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \psi_n \right\| \cdot 2 + \left\| \mathcal{O} \Lambda_{A,+} \frac{1}{H_0 + \mu} \psi_n \right\|^2. \end{aligned} \quad (5.13)$$

Since  $D_A \Lambda_{A,+} (H_0 + \mu)^{-1} = \Lambda_{A,+} D_A \Lambda_{A,+} (\Lambda_{A,+} D_A \Lambda_{A,+} + \Lambda_{A,+} E_f \Lambda_{A,+} + \mu)^{-1} \leq 1$ , all terms on the rhs of (5.13) are bounded.

Concerning the last term of (5.10), we will establish the compactness of the operator  $K := \chi_0 \Lambda_{A,+} (H_0 + \mu)^{-1}$ . Then  $\| \tilde{\psi}_n \| \rightarrow 0$  for  $n \rightarrow \infty$ . Collecting results, this shows that the second term of (5.8) can be made smaller than  $\epsilon/2$  for  $n$  sufficiently large and thus proves the desired compactness of the operator (5.4).

The strategy to show the compactness of  $K$  is to start with the operator  $K_1 := \chi_0 (p^2 + m^2)^{-\frac{1}{2}}$  which is compact as a product of bounded functions  $f(\mathbf{x})$ ,  $g(\mathbf{p})$ , each of which tending to zero as  $x$ , respectively  $p$ , go to infinity (see, e.g., [31, lemma 7.10]). Then, bounded operators  $\mathcal{O}_1, \mathcal{O}_2$  are constructed such that  $K_1 \cdot \prod \mathcal{O}_i = K$ .

Let  $\mathcal{O}_1 := \sqrt{p^2 + m^2} D_A^{-1}$ . For showing the boundedness of  $\mathcal{O}_1$  let  $\psi := D_A^{-1} \varphi$ . Then from the diamagnetic inequality and (4.6),

$$\begin{aligned} \| \mathcal{O}_1 \varphi \|^2 &= (\psi, (p^2 + m^2) \psi) \leq (\psi, (E_A^2 + e|\mathbf{B}|) \psi) \\ &\leq \left( 1 + \frac{\kappa e}{1 - \kappa e} \right) \| \varphi \|^2 + \frac{e C_\kappa}{1 - \kappa e} \| D_A^{-2} \| \| \varphi \|^2, \end{aligned} \quad (5.14)$$

the rhs being obviously bounded.

With  $\mathcal{O}_2 := D_A \Lambda_{A,+} (H_0 + \mu)^{-1} \leq 1$  (as shown above), we arrive at  $K_1 \cdot \mathcal{O}_1 \cdot \mathcal{O}_2 = K$ . □

We remark that in the same way the compactness of  $x^{-\frac{1}{2}} (|D_A| + \mu)^{-1}$  on  $L_2(\mathbb{R}^3) \otimes \mathbb{C}^4$  can be shown. The only additional ingredient is the boundedness of  $D_A (|D_A| + \mu)^{-1}$  in the equation corresponding to (5.13), which follows from  $\| D_A (|D_A| + \mu)^{-1} \psi_n \|^2 = (\psi_n, |D_A|^2 (|D_A| + \mu)^{-2} \psi_n) \leq \| \psi_n \|^2$ .

### 5.2. Boundedness of $(H_0 + \mu)^\lambda (H^{(2)} + \mu)^{-1}$

Let first  $\lambda = 1$ . From (3.7) and (3.21), we have the relative form boundedness of the potential for  $\psi \in H_1(\mathbb{R}^3) \otimes \mathbb{C}^4$  and  $\psi_+ := \Lambda_{A,+} \psi$ ,

$$\| \Lambda_{A,+} (V + B_{2m}) \Lambda_{A,+} \psi \|^2 \leq \| V \psi_+ \|^2 + \| B_{2m} \psi_+ \|^2 \leq a_0 \| D_A \psi_+ \|^2 + b_0 \| \psi_+ \|^2 \quad (5.15)$$

with

$$a_0 := \frac{\gamma}{\gamma_1} + \frac{\gamma^2}{2\gamma_1} \left( \frac{1}{\gamma_1} + \frac{1}{\gamma_0} + \frac{c_B}{m} \right), \quad b_0 := \gamma d_B + C, \quad (5.16)$$

where  $c_B$  is defined in (3.6). We have to restrict  $\gamma < \tilde{\gamma}_c$  such that  $a_0 < 1$ .  $\tilde{\gamma}_c$  depends on  $\mathbf{B}$ , its maximum value (for  $\mathbf{B} = 0$ ) being  $\tilde{\gamma}_c^{(0)} = 0.319$  ( $Z \leq 43$ ), obtained as solution to  $2\gamma + \gamma^2 (2 + \frac{\pi}{2}) = 1$ .

Let  $\epsilon := 1 - a_0$  with  $0 < \epsilon < 1$ . With  $\psi := (H^{(2)} + \mu)^{-1}\psi_+$ , we want to show

$$\left\| (H_0 + \mu) \frac{1}{H^{(2)} + \mu} \psi_+ \right\|^2 = \|(H_0 + \mu)\psi\|^2 \stackrel{!}{\leq} c_1^2 \|\psi_+\|^2 = c_1^2 \|(H^{(2)} + \mu)\psi\|^2 \tag{5.17}$$

for a suitable  $c_1 > 0$ . We estimate, using  $\|D_A \psi_+\| = \|\Lambda_{A,+} D_A \Lambda_{A,+} \psi\| \leq \|H_0 \psi\|$  and  $\|\psi_+\| \leq \|\Lambda_{A,+}\| \|\psi\|$ ,

$$\begin{aligned} c_1 \|(H^{(2)} + \mu)\psi\| &\geq c_1 \|(H_0 + \mu)\psi\| - c_1 \|\Lambda_{A,+}(V + B_{2m})\Lambda_{A,+}\psi\| \\ &\geq c_1 \|(H_0 + \mu)\psi\| - c_1 \{a_0 \|H_0 \psi\| + b_0 \|\psi\|\} \\ &\stackrel{!}{\geq} c_1 \|(H_0 + \mu)\psi\| + (1 - c_1)(\|H_0 \psi\| + \mu \|\psi\|) \geq \|(H_0 + \mu)\psi\|. \end{aligned} \tag{5.18}$$

Condition (5.18) is satisfied if  $-c_1 a_0 \geq 1 - c_1$  as well as  $-c_1 b_0 \geq (1 - c_1)\mu$ , requiring the choice  $c_1 \geq 1/\epsilon$  and  $\mu \geq c_1 b_0 / (c_1 - 1)$ .

For  $\lambda = \frac{1}{2}$ , the bound on  $\gamma$  can be improved by working with quadratic forms. From (3.22), we have

$$(\psi, \Lambda_{A,+}(V + B_{2m})\Lambda_{A,+}\psi) \geq -a_1 (\psi, \Lambda_{A,+} D_A \Lambda_{A,+} \psi) - \tilde{C}(\psi, \psi) \tag{5.19}$$

with  $a_1 := a_0 - \gamma(\frac{1}{\gamma_1} - \frac{1}{\gamma_0})$ . Trivially, we have  $(H_0 + \mu)^{\frac{1}{2}}(H^{(2)} + \mu)^{-1} = (H_0 + \mu)^{\frac{1}{2}}(H^{(2)} + \mu)^{-\frac{1}{2}} \cdot (H^{(2)} + \mu)^{-\frac{1}{2}}$  where the last factor is bounded. For the boundedness of the other factor, we use the strategy of (5.17) to require  $\|(H_0 + \mu)^{\frac{1}{2}}\psi\|^2 \leq c_2 \|(H^{(2)} + \mu)^{\frac{1}{2}}\psi\|^2$  which is satisfied if  $c_2 \geq 1/(1 - a_1)$  and  $\mu \geq c_2 \tilde{C}/(c_2 - 1)$ . The necessary condition for  $c_2 > 0$  is  $a_1 < 1$ , i.e. inequality (4.3). The corresponding maximum value for  $\gamma$  is  $\gamma_c^{(0)} = 0.353$ .

### 5.3. Compactness of $R_\mu(V)$

We take  $\lambda = \frac{1}{2}$  and decompose

$$-\frac{1}{H_0 + \mu} \Lambda_{A,+} V \Lambda_{A,+} \frac{1}{(H_0 + \mu)^{\frac{1}{2}}} = \gamma \left\{ \frac{1}{H_0 + \mu} \Lambda_{A,+} \frac{1}{x^{\frac{1}{2}}} \right\} \left[ \frac{1}{x^{\frac{1}{2}}} \Lambda_{A,+} \frac{1}{(H_0 + \mu)^{\frac{1}{2}}} \right]. \tag{5.20}$$

The factor in square brackets is bounded according to a (5.5)-type estimate by using that  $\Lambda_{A,+}|D_A|\Lambda_{A,+} \leq H_0 + \mu$ . Together with lemma 2 and the result of section 5.2, this proves the compactness of  $R_\mu(V)$  for  $\gamma < \gamma_c$  determined from (4.3).

### 5.4. Compactness of $R_\mu(B_{2m})$

According to the four contributions of  $B_{2m}$  from (2.12), we define

$$\frac{1}{H_0 + \mu} (\Lambda_{A,+} B_{2m} \Lambda_{A,+}) \frac{1}{(H_0 + \mu)^\lambda} =: -\frac{\gamma}{4} \sum_{i=1}^4 \mathcal{O}_i(\lambda). \tag{5.21}$$

For  $i = 1$ , we take  $\lambda = 1$  and decompose

$$\mathcal{O}_1(1) = \left\{ \frac{1}{H_0 + \mu} \Lambda_{A,+} \frac{1}{x^{\frac{1}{2}}} \right\} \cdot \left[ \frac{1}{x^{\frac{1}{2}}} F_A \tilde{D}_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \right]. \tag{5.22}$$

In order to show the boundedness of the operator in square brackets, we use a (5.5)-type estimate for  $x^{-\frac{1}{2}}$  and note that  $F_A \tilde{D}_A \Lambda_{A,+} (H_0 + \mu)^{-1}$  is bounded. It then remains to show the boundedness of  $M := |D_A|^{\frac{1}{2}} F_A \tilde{D}_A \Lambda_{A,+} (H_0 + \mu)^{-1}$ . We estimate for  $\varphi \in L_2(\mathbb{R}^3) \otimes \mathbb{C}^4$ , noting that  $F_A \tilde{D}_A = -\tilde{D}_A F_A$ ,

$$\begin{aligned} \|M\varphi\|^2 &= \left( \tilde{D}_A F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi, |D_A| \tilde{D}_A F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right) \\ &\leq \left\| \tilde{D}_A F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \right\| \|\varphi\| \cdot \left\| D_A F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\|. \end{aligned} \tag{5.23}$$

We use (3.16) to commute  $D_A$  with  $F_A$ , being left with two terms involving the potential  $V$ . In turn, these terms can be estimated according to (3.7) by replacing  $V$  with  $|D_A|$  plus a bounded remainder. For example, we get

$$\begin{aligned} \left\| \frac{1}{2} \tilde{D}_A V \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\| &\leq \frac{\gamma}{2\gamma_1} \|\tilde{D}_A\| \left\| |D_A| \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\| \\ &\quad + \frac{\gamma d_B}{2} \|\tilde{D}_A\| \left\| \Lambda_{A,+} \frac{1}{H_0 + \mu} \right\| \|\varphi\| \end{aligned} \tag{5.24}$$

which obviously is bounded.

For  $i = 2$ , we take  $\lambda = \frac{1}{2}$  and decompose

$$\mathcal{O}_2\left(\frac{1}{2}\right) = \frac{1}{H_0 + \mu} \Lambda_{A,+} \tilde{D}_A F_A (|D_A| + \mu) \cdot \left\{ \frac{1}{|D_A| + \mu} \frac{1}{x^{\frac{1}{2}}} \right\} \cdot \left[ \frac{1}{x^{\frac{1}{2}}} \Lambda_{A,+} \frac{1}{(H_0 + \mu)^{\frac{1}{2}}} \right]. \tag{5.25}$$

Referring to our previous considerations, it remains to show the boundedness of the adjoint of the first term,  $|D_A| F_A \tilde{D}_A \Lambda_{A,+} (H_0 + \mu)^{-1}$ , since  $\mu(H_0 + \mu)^{-1} \Lambda_{A,+} \tilde{D}_A F_A$  is trivially bounded (and since any bounded operator has a bounded adjoint). With  $|D_A| F_A \tilde{D}_A = -D_A F_A$ , we arrive at the last term of (5.23), the boundedness of which has just been shown.

For  $i = 3$ , we take again  $\lambda = 1$ . Then,

$$\begin{aligned} \mathcal{O}_3(1) &= \frac{1}{H_0 + \mu} \Lambda_{A,+} \tilde{D}_A \frac{1}{x} F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \\ &= \tilde{D}_A \cdot \left\{ \frac{1}{H_0 + \mu} \Lambda_{A,+} \frac{1}{x^{\frac{1}{2}}} \right\} \cdot \left[ \frac{1}{x^{\frac{1}{2}}} F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \right], \end{aligned} \tag{5.26}$$

of which the first factor is compact and the second factor bounded. For the factor in square brackets we estimate according to (5.5), and further

$$\left\| |D_A|^{\frac{1}{2}} F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\|^2 \leq \left\| F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\| \cdot \left\| |D_A| F_A \Lambda_{A,+} \frac{1}{H_0 + \mu} \varphi \right\|. \tag{5.27}$$

Since  $\| |D_A| \tilde{\varphi} \|^2 = (\tilde{\varphi}, D_A^2 \tilde{\varphi}) = \| D_A \tilde{\varphi} \|^2$ , the second factor agrees with the one from (5.23).

For  $i = 4$  and  $\lambda = 1$ , we have  $\mathcal{O}_4(1) = (H_0 + \mu)^{-1} \Lambda_{A,+} F_A \frac{1}{x} \tilde{D}_A \Lambda_{A,+} (H_0 + \mu)^{-1} = \mathcal{O}_3(1)^*$ . Together with the result from section 5.2, this proves compactness of  $R_\mu(B_{2m})$  for  $\gamma < \tilde{\gamma}_c$  defined below (5.16).

### 6. The essential spectrum

For the Schrödinger operator with purely magnetic field,  $\mathbf{p}_A^2$ , it was shown, following the work of Jörgens [17], that its essential spectrum is given by

$$\sigma_{\text{ess}}(\mathbf{p}_A^2) = [0, \infty) \tag{6.1}$$

provided  $\mathbf{A} \in L_{2,\text{loc}}(\mathbb{R}^3)$  and

$$N_A(\mathbf{x}) = \int_{|\mathbf{x}-\mathbf{y}|\leq 1} |\mathbf{A}(\mathbf{y})|^2 \, d\mathbf{y} \longrightarrow 0 \tag{6.2}$$

as  $\mathbf{x} \rightarrow \infty$  ([20], see also [33]). In particular, condition (6.2) is satisfied if  $\mathbf{B} \rightarrow 0$  as  $\mathbf{x} \rightarrow \infty$  [20]. It is, however, easy to show that it is sufficient that  $N_B(\mathbf{x}) \rightarrow 0$  (as  $\mathbf{x} \rightarrow \infty$ ) for (6.2) to hold. We use the relation between  $\mathbf{A}$  and  $\mathbf{B}$  introduced in [11],

$$\mathbf{A}(\mathbf{y}) = \int_0^1 t \, dt \, \mathbf{B}(\mathbf{x} + t(\mathbf{y} - \mathbf{x})) \wedge (\mathbf{y} - \mathbf{x}), \tag{6.3}$$

which satisfies  $\nabla \times \mathbf{A} = \mathbf{B}$  (since  $\nabla \cdot \mathbf{B} = 0$ ). Then we have, substituting  $\mathbf{z} := \mathbf{y} - \mathbf{x}$ ,

$$\begin{aligned} N_A(\mathbf{x}) &= \int_{z \leq 1} |\mathbf{A}(\mathbf{z} + \mathbf{x})|^2 \, d\mathbf{z} \\ &= \int_{z \leq 1} d\mathbf{z} \int_0^1 t \, dt \int_0^1 \tau \, d\tau (\mathbf{B}(\mathbf{x} + t\mathbf{z}) \wedge \mathbf{z})(\mathbf{B}(\mathbf{x} + \tau\mathbf{z}) \wedge \mathbf{z}). \end{aligned} \tag{6.4}$$

We estimate  $|\mathbf{B} \wedge \mathbf{z}| \leq |\mathbf{B}|$  (since  $z \leq 1$ ) and factorize the integrand according to  $\frac{t^{1+\epsilon}}{\tau^\epsilon} |\mathbf{B}(\mathbf{x} + t\mathbf{z})| \cdot \frac{\tau^{1+\epsilon}}{t^\epsilon} |\mathbf{B}(\mathbf{x} + \tau\mathbf{z})|$  with, e.g.,  $\epsilon = \frac{1}{4}$ . Applying the Schwarz inequality, we get upon substituting  $\boldsymbol{\xi} := t\mathbf{z}$  for  $\mathbf{z}$

$$\begin{aligned} N_A(\mathbf{x}) &\leq \left( \int_0^1 \frac{d\tau}{\tau^{2\epsilon}} \right) \int_0^1 t^{2+2\epsilon} \, dt \int_{z \leq 1} d\mathbf{z} |\mathbf{B}(\mathbf{x} + t\mathbf{z})|^2 \\ &= \int_0^1 \frac{d\tau}{\tau^{\frac{1}{2}}} \int_0^1 \frac{dt}{t^{\frac{1}{2}}} \int_{\xi \leq t} d\boldsymbol{\xi} |\mathbf{B}(\mathbf{x} + \boldsymbol{\xi})|^2 \leq 4 \int_{\xi \leq 1} d\boldsymbol{\xi} |\mathbf{B}(\mathbf{x} + \boldsymbol{\xi})|^2 \end{aligned} \tag{6.5}$$

which, upon assumption, tends to 0 as  $\mathbf{x} \rightarrow \infty$ .

A further consequence of  $N_B(\mathbf{x}) \rightarrow 0$  (as  $\mathbf{x} \rightarrow \infty$ ) is that  $e\sigma\mathbf{B}$  is  $\mathbf{p}_A^2$ -compact [32, theorem 5.2.2]. Thus, the essential spectrum of the Pauli operator  $(\sigma\mathbf{p}_A)^2$  is also given by  $[0, \infty)$ . Accordingly,  $\sigma_{\text{ess}}(D_A^2) = [m^2, \infty)$ , and therefore

$$\sigma_{\text{ess}}(E_A) = [m, \infty). \tag{6.6}$$

In fact, let  $\lambda^2 \in [m^2, \infty)$  and  $\lambda > 0$ . Then, there exists a normalized sequence  $\varphi_n \in C_0^\infty(\mathbb{R}^3) \otimes \mathbb{C}^2$  with  $\varphi_n \xrightarrow{w} 0$  such that  $\|(E_A - \lambda)(E_A + \lambda)\varphi_n\| \rightarrow 0$  as  $n \rightarrow \infty$ . Let  $\phi \in C_0^\infty(\mathbb{R}^3) \otimes \mathbb{C}^2$  and note that  $C_0^\infty \subset H_2 \subset H_1 \subset L_2$ . Then,  $(E_A + \lambda)\phi \in H_1(\mathbb{R}^3) \otimes \mathbb{C}^2 = \mathcal{D}(E_A)$  and  $(\phi, (E_A + \lambda)\varphi_n) = ((E_A + \lambda)\phi, \varphi_n) \rightarrow 0$  (since  $\varphi_n \xrightarrow{w} 0$ ) as  $n \rightarrow \infty$ . Moreover,  $\liminf_{n \rightarrow \infty} \|(E_A + \lambda)\varphi_n\| \geq \liminf_{n \rightarrow \infty} \|(m + \lambda)\varphi_n\| = m + \lambda > 0$ , which shows that  $\tilde{\varphi} := (E_A + \lambda)\varphi_n \xrightarrow{w} 0$ , such that  $\lambda \in \sigma_{\text{ess}}(E_A)$  [34, theorem 7.24, p 191].

In order to derive  $\sigma_{\text{ess}}(D_A)$  from (6.6), we note that  $\sigma_{\text{ess}}(-E_A) = (-\infty, -m]$ . Moreover, since a unitary transformation does not change the essential spectrum, we have from (1.7)

$$\begin{aligned} \sigma_{\text{ess}}(D_A) &= \sigma_{\text{ess}}(U_0 D_A U_0^{-1}) = \sigma_{\text{ess}}(\beta E_A) \\ &= \sigma_{\text{ess}} \begin{pmatrix} E_A & \\ & 0 \end{pmatrix} \cup \sigma_{\text{ess}} \begin{pmatrix} 0 & \\ & -E_A \end{pmatrix} = [m, \infty) \cup (-\infty, -m]. \end{aligned} \tag{6.7}$$

It was proven earlier [11, theorem 1.4] that  $\sigma_{\text{ess}}(D_A) = (-\infty, -m] \cup [m, \infty)$  under somewhat stronger assumptions (e.g.,  $\mathbf{B}(\mathbf{x}) \rightarrow 0$  as  $\mathbf{x} \rightarrow \infty$ ), the proof being similar to the one given in [5, p 117] for the Schrödinger case.

From the decomposition of  $D_A$  into its (disjoint) positive and negative part,  $D_A = \Lambda_{A,+} D_A \Lambda_{A,+} + \Lambda_{A,-} D_A \Lambda_{A,-}$ , we get  $\sigma_{\text{ess}}(\Lambda_{A,+} D_A \Lambda_{A,+}) = \sigma_{\text{ess}}(E_A) = [m, \infty)$ .

Together with theorem 2 we have thus proven

**Theorem 3.** *Let  $H^{(2)}$  be the ‘magnetic’ Jansen–Hess operator, let the vector potential  $\mathbf{A} \in L_{2,\text{loc}}(\mathbb{R}^3)$ , let the magnetic field obey  $N_B(\mathbf{x}) \rightarrow 0$  for  $\mathbf{x} \rightarrow \infty$  with finite field energy  $E_f$ . Then for a Coulomb potential with strength  $\gamma < \tilde{\gamma}_c$ , the essential spectrum is given by*

$$\sigma_{\text{ess}}(H^{(2)}) = [m, \infty) + E_f, \tag{6.8}$$

where  $\tilde{\gamma}_c$  is defined in theorem 2.



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